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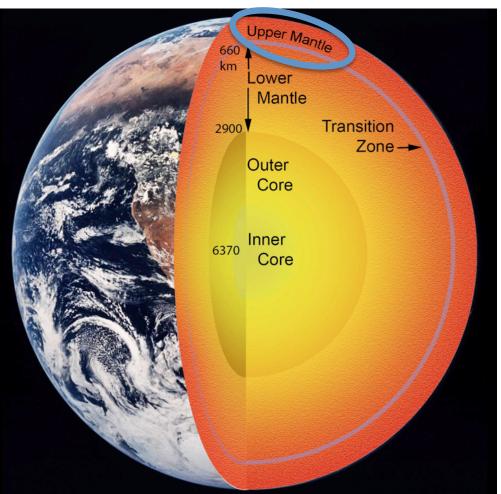


Roberta Rudnick (U Maryland, USA), Lillian Schaffer (U Houston, USA), Jon Snow (U Houston, USA), Larry Taylor (U Tennessee) Alan Woodland (U Frankfurt, Germany)

- ♦ (Water in the exosphere)
- ♦ Water in the crust
- **♦ Water in the mantle**
- ♦ Water in the core



Glacier in the Miller range, Antarctica



*Image Credit: Japan Times* 

- ♦ Type of data
- ♦ Water in the mantle
  - ♦ Definition
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    - ♦ H diffusion
    - ♦ Water in the continental mantle lithosphere
    - ♦ Water in the lithosphere
    - ♦ Water in the oceanic lithosphere
- ♦ The big picture
  - ♦ Water in the Earth layers
  - ♦ Fluxes
- ♦ Comparison with other differentiated planetary bodies

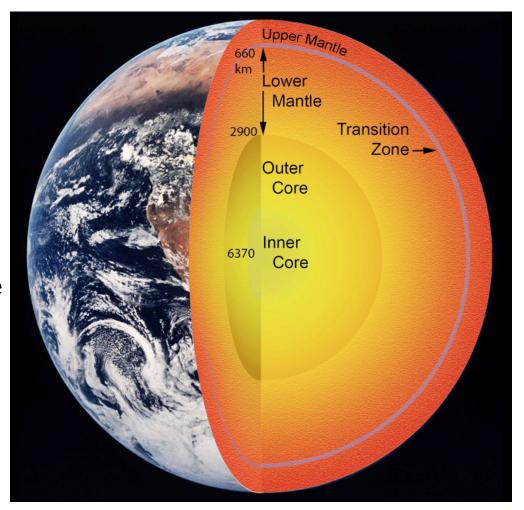


Image Credit: Japan Times

## Type of data for water in the deep Earth

#### ♦ Crust

♦ Direct samples: crustal rocks

#### Mantle lithosphere

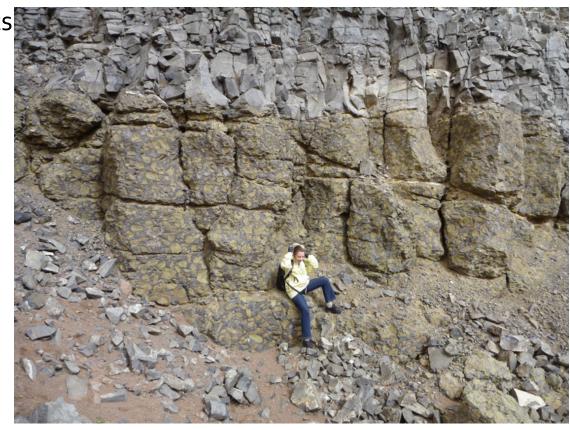
♦ Direct samples: Peridotites, pyroxenites samples

#### **♦** Asthenosphere

- ♦ Direct samples: Rare deep diamond inclusions
- ♦ HP Experiments
- ♦ First principle calculations

#### **♦Core**

- ♦ H solubility in Fe-Ni metal
- ♦ Accretion models



Mantle xenoliths in basaltic flow, Jieshaba, China

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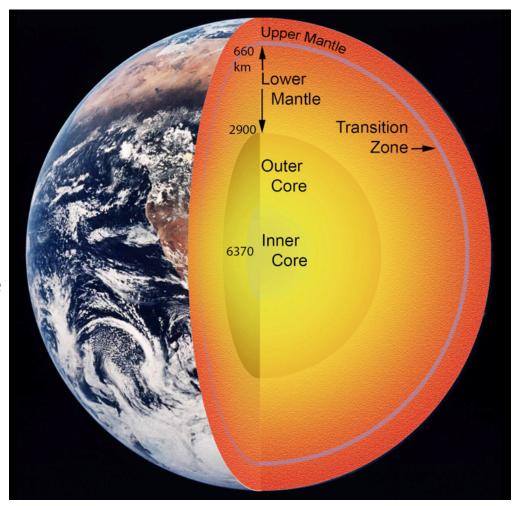
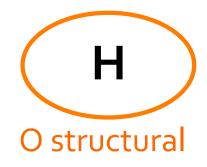


Image Credit: Japan Times

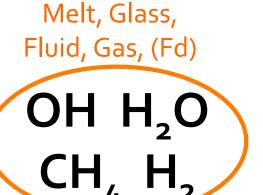
#### Disclaimer about "water" in the mantle

"Water" in anhydrous minerals = Hydrogen

Ol, Px, Gr, Fd, Apatite, Amph, Mica, Ring, Wads, Perov Phase B, D, H and Egg







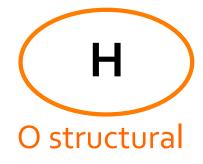
♦ Calculated in ppm wt H<sub>2</sub>O (<1-1000')</p>

#### Disclaimer about "water" in the mantle

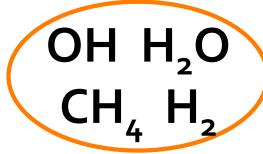
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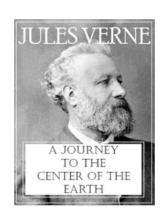
Melt, Glass, Fluid, Gas, (Fd)

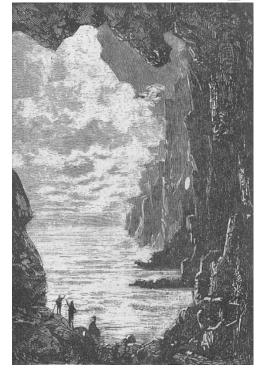






♦ Calculated in ppm wt H<sub>2</sub>O (<1-1000')</p>





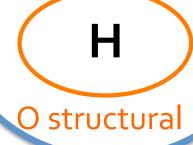
Equivalent of several Earth's oceans in the mantle

#### Disclaimer about "water" in the mantle

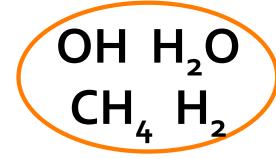
♦ "Water" in anhydrous minerals = Hydrogen

Ol, Px, Gr, Fd, Ring, Wads, Perov

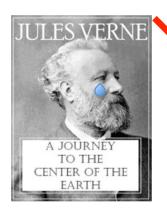
Apatite, Amph, Mica, Phase B, D, H and Egg Melt, Glass, Fluid, Gas, (Fd)



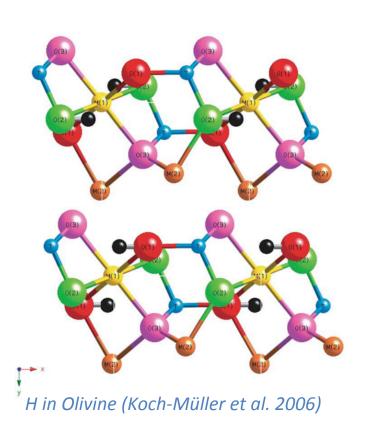




 $\Leftrightarrow$  Calculated in ppm wt H<sub>2</sub>O (<1-1000')





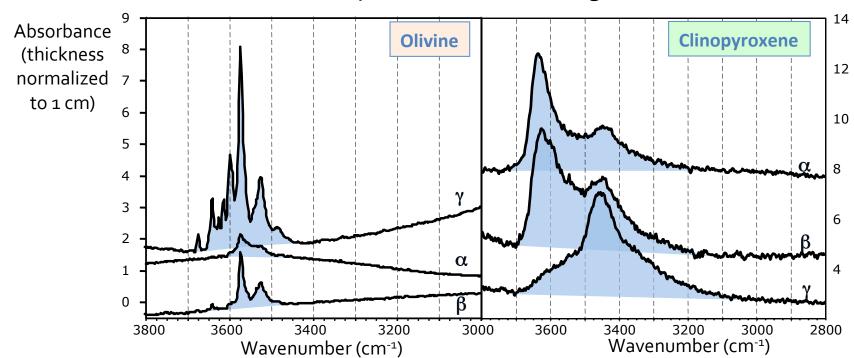


# **Techniques**

- ♦ FTIR → H2O content
  - → speciation
  - → location H mineral defects
- ♦ SIMS → H<sub>2</sub>O content
  - → H isotopes
  - $\rightarrow$  Cl, F

Detection limits FTIR <0.5 ppm H2O SIMS ≤ 2 ppm H2O

FTIR spectra in the O-H region



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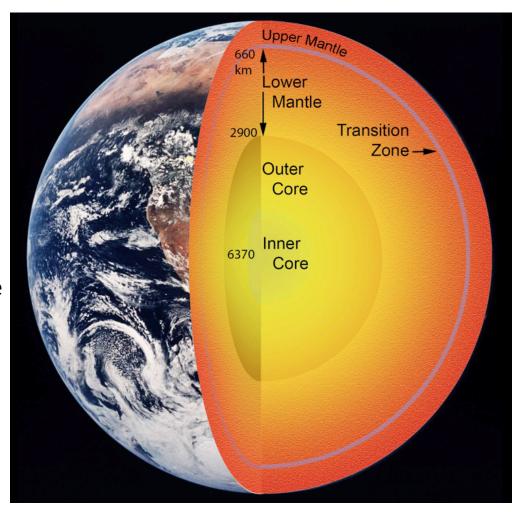


Image Credit: Japan Times

# Why is water important?

#### ♦ Water lowers the solidus of mantle lithologies:

- Facilitates partial melting, lower T

#### ♦ Influence on magmatism

- Magma composition
- Magma transfer and eruption style
  - → origin of oceans & atmosphere

#### ♦ Water and rheology:

- Presence of water in olivine makes it weaker
  - → crucial for plate tectonics
- Melt circulation and eruption style

#### ♦ Water and remote sensing of the deep Earth:

- Seismic properties: seismic wave attenuation & anisotropy
- Electrical conductivity
- Thermal conductivity



Image Credit: J. Head Brown U., Stromboli

Gaetani & Grove, 1998; Green, 1973; Hirose & Kawamoto, 1995; Chopra & Paterson, 1984; Dixon et al., 2004; Drury, 1991; Hirth & Kohlstedt, 1996; Hirth et al., 2000; Justice et al., 1982; Karato, 1993; 2004; 2006; 2010; Mackwell et al., 1985; Mei & Kohlstedt, 2000; Walker et al., 2007; Jung & Karato, 2001; Demouchy et al 2012; Jones, Evans, Muller, Fullea, Pommier 1990-2015; Lizzaralde et al 1995; Sarafian et al. 2015; Udata et al 2003; Hofmeister 2004

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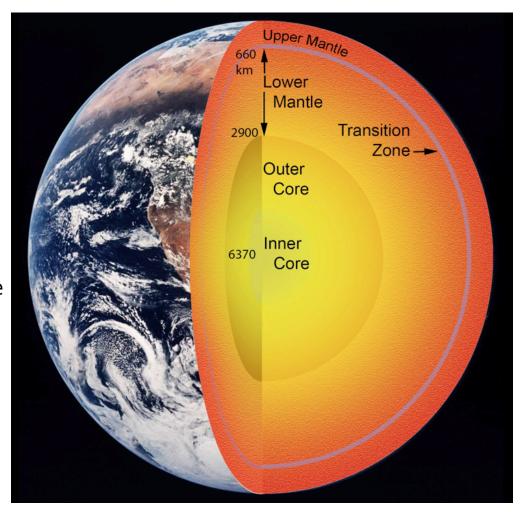


Image Credit: Japan Times

# H loss during xenolith ascent in host magma?

♦ H diffuses quickly through OI & Px

(e.g. Mackwell & Kohlstedt 1990; Ingrin et al 1995-2006, Stalder & Skogby 2003)

♦ Px have homogeneous water contents

♦ Ol can record H loss

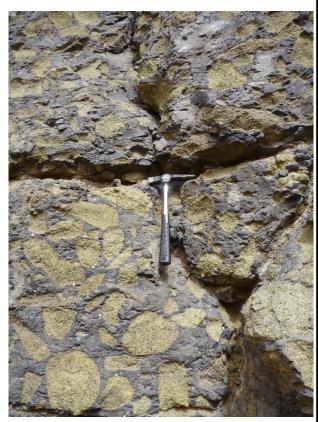
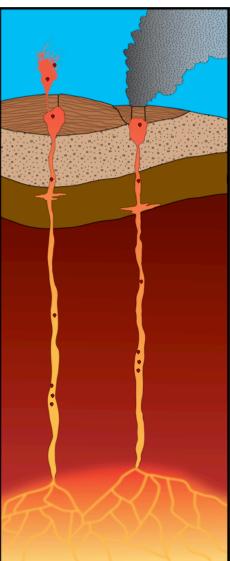


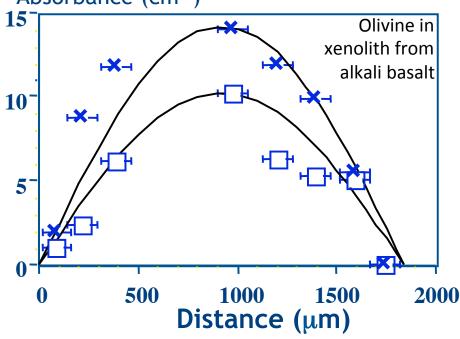
Figure stolen from AGU Chapman conference website (xenoliths added)



D E F

Olivine from Mexican mantle xenolith

Absorbance (cm<sup>-2</sup>)



# H loss during xenolith ascent in host magma?

- H diffuses quickly through Ol & Px at magmatic temperatures
  - (e.g. Mackwell & Kohlstedt 1990; Ingrin et al 1995, 1999, Stalder & Skogby 2003)
- But H not always lost because H diffusion is coupled with that of slower elements (e.g. Berry et al 2007, Novella et al 2015, Skogby, Stalder, Sundval 1989-2009)
- Check:

20

100

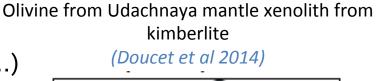
With diffusion modeling

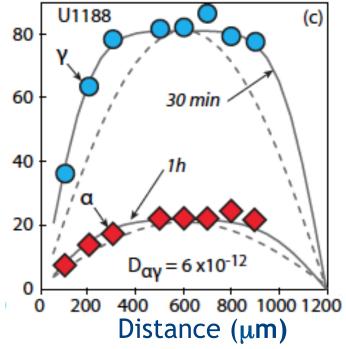
200

Distance (µm)

300

If co-var with other elements (Al, Ti, REE...) Absorbance (cm<sup>-2</sup>) 80 **60**. Olivine in xenolith from Kimberlite (Kaapvaal) 40 Peslier et al 2008





→ Data shown here are mantle [H<sub>2</sub>O] values

400

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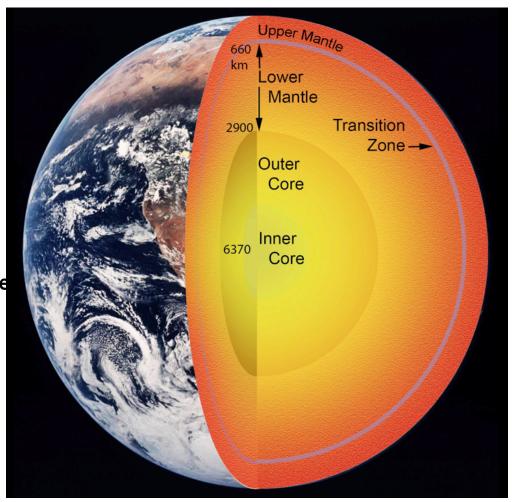
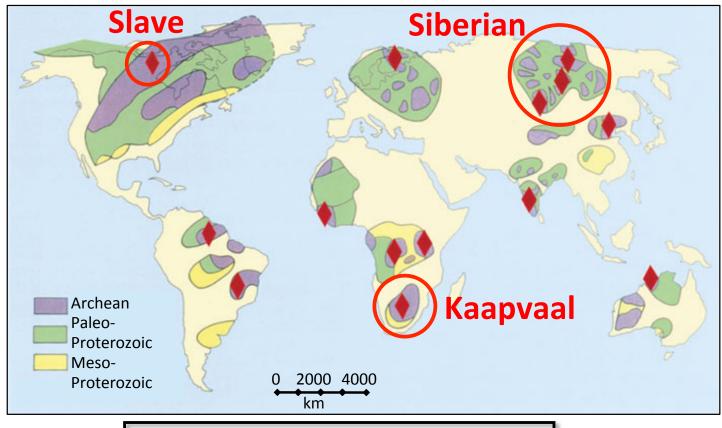
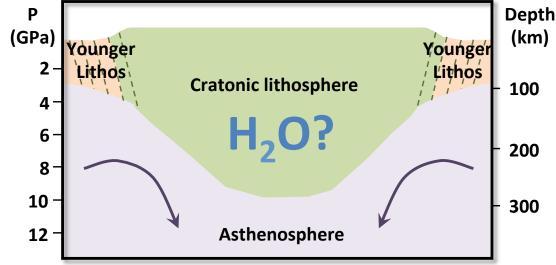
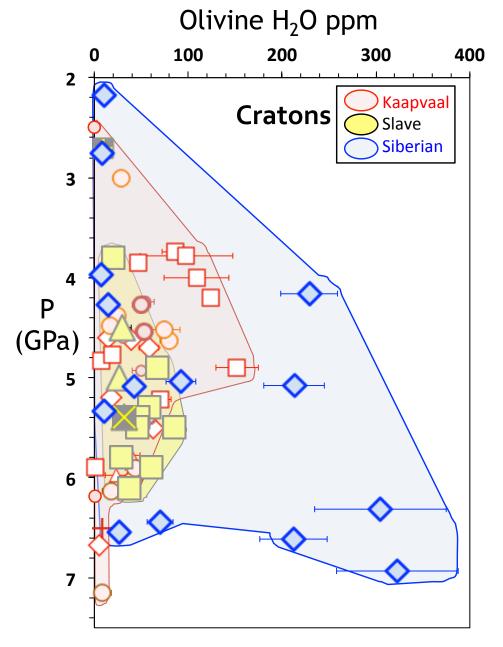


Image Credit: Japan Times

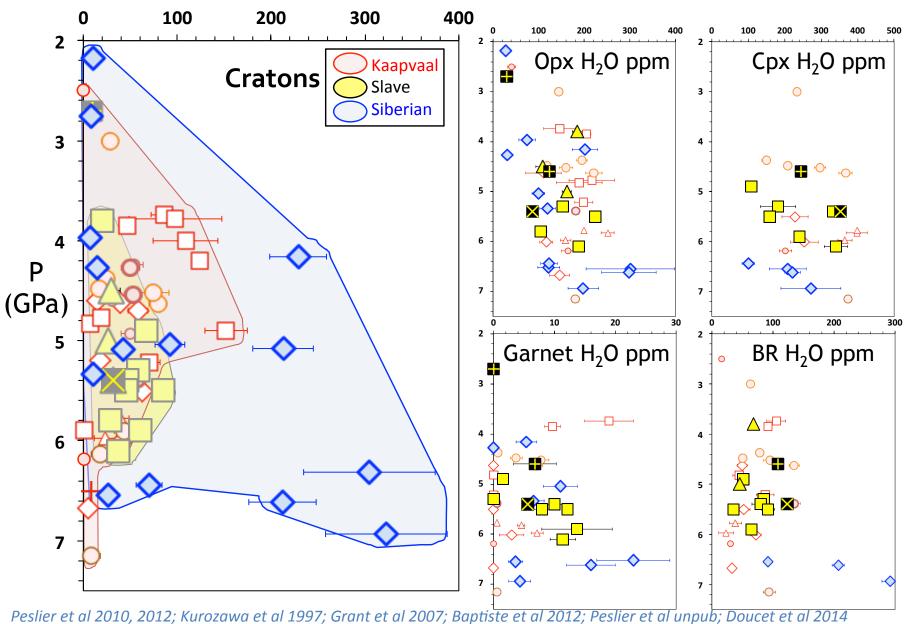
♦ Cratons





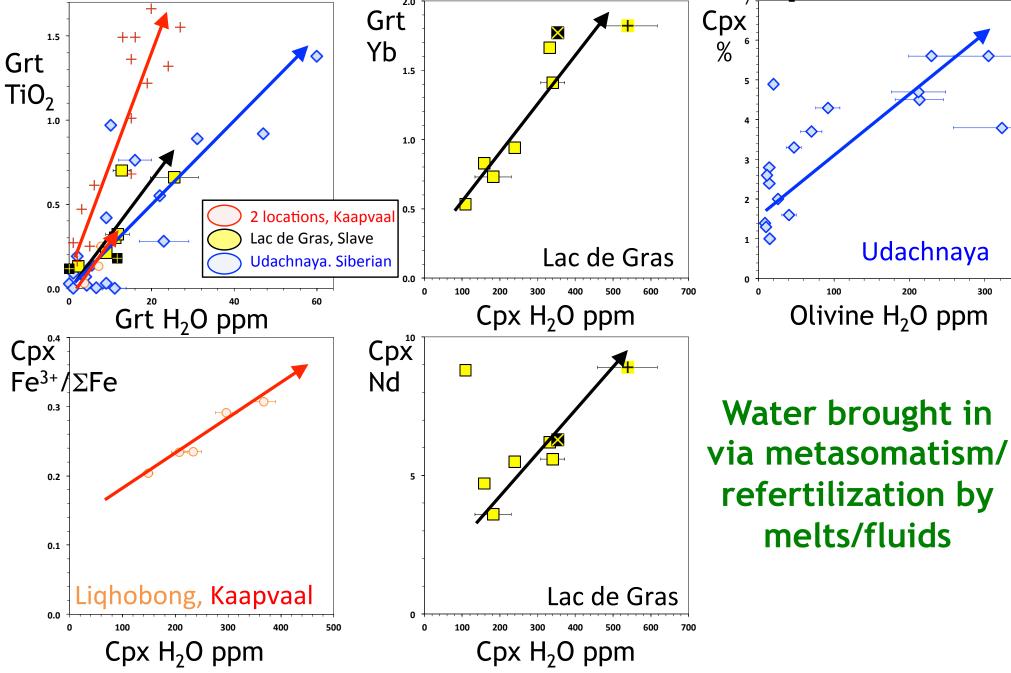


Olivine H<sub>2</sub>O ppm



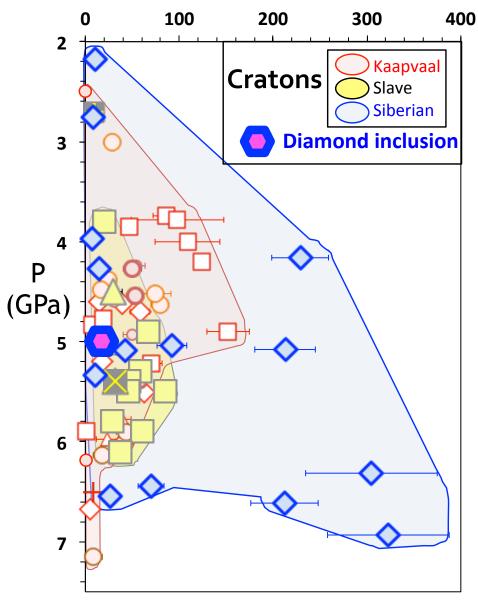
Olivine H<sub>2</sub>O ppm 100 200 300 400 0 100 Opx H<sub>2</sub>O ppm Cpx H<sub>2</sub>O ppm Kaapvaal **Cratons** Slave Siberian P (GPa) **Diamond inclusion** Garnet H<sub>2</sub>O ppm 6 7 1 mm

Peslier et al 2010, 2012; Kurozawa et al 1997; Grant et al 2007; Baptiste et al 2012; Peslier et al unpub; Doucet et al 2014, Novella et al 2015, Taylor et al 2016

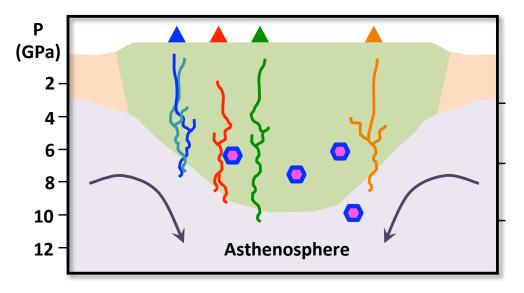


Bell et al., 1992; Peslier et al 2012; Doucet et al 2014; Peslier et al unpublished

Olivine H<sub>2</sub>O ppm



Xenoliths NOT representative of whole mantle lithosphere



- Over-representation of metasomatized water-rich peridotite near melt/fluid channels
- Overall mantle lithos dry (preserved in diamond mineral inclusions)
- > Water has a role in cratonic root long term longevity

Peslier et al 2010, 2012; Kurozawa et al 1997; Grant et al 2007; Baptiste et al 2012; Peslier et al unpub; Doucet et al 2014, Novella et al 2015, Taylor et al 2016

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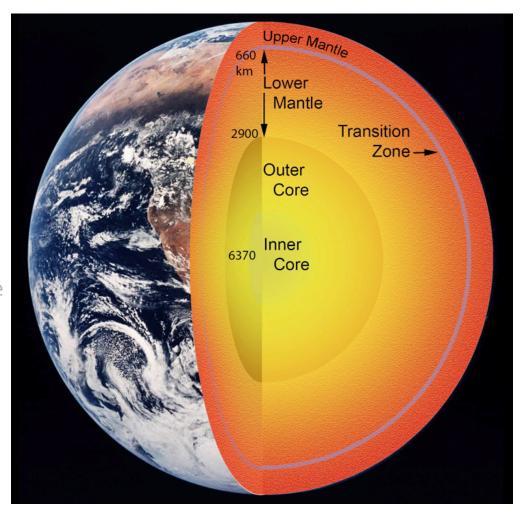
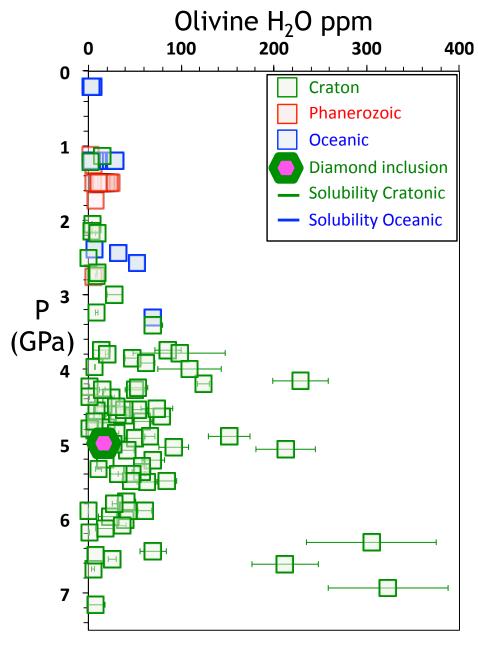


Image Credit: Japan Times



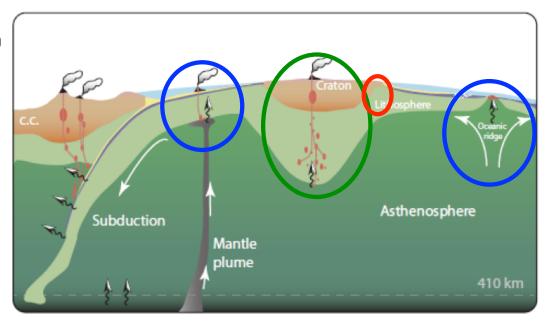
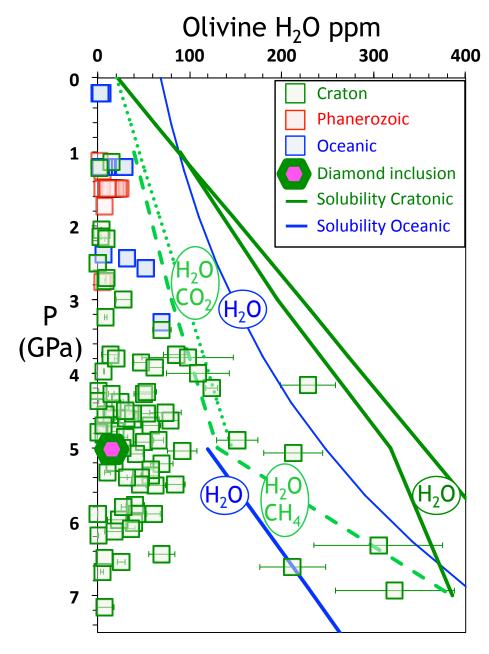


Image Credit: Demouchy & Bolfan-Casanova 2016

Off-craton olivine: most have lost H during xeno ascent

Peslier et al 2010- 2015-Unpub; Kurozawa et al 1997; Grant et al 2007; Baptiste et al 2012; Doucet et al 2014, Novella et al 2015, Taylor et al 2016; Warren & Hauri 2014



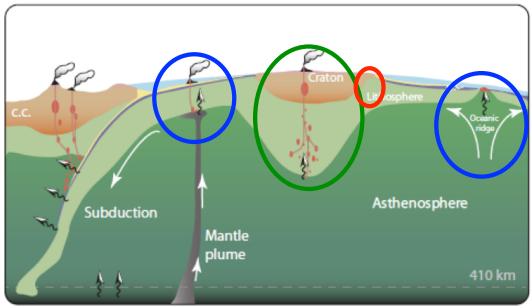
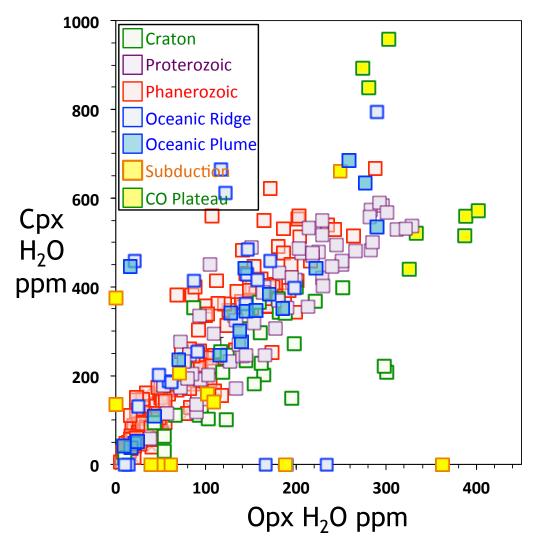


Image Credit: Demouchy & Bolfan-Casanova 2016

- Off-craton olivine: most have lost H during xeno ascent
- Max [H2O] in olivine limited by solubility?
- Max [H2O] in olivine depends on local melt/fluid water activity
- Mantle lithosphere is unsaturated in water

Peslier et al 2010- 2015-Unpub; Kurozawa et al 1997; Grant et al 2007; Baptiste et al 2012; Doucet et al 2014, Novella et al 2015, Taylor et al 2016; Warren & Hauri 2014; Férot & Bolfan-Casanova 2012, Tenner, Ardia et al 2012; Yang et al 2014-15; Demouchy & Bolfan-Casanoca 2016



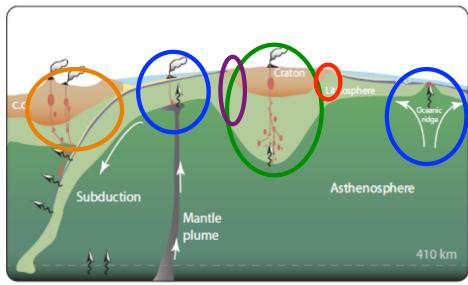
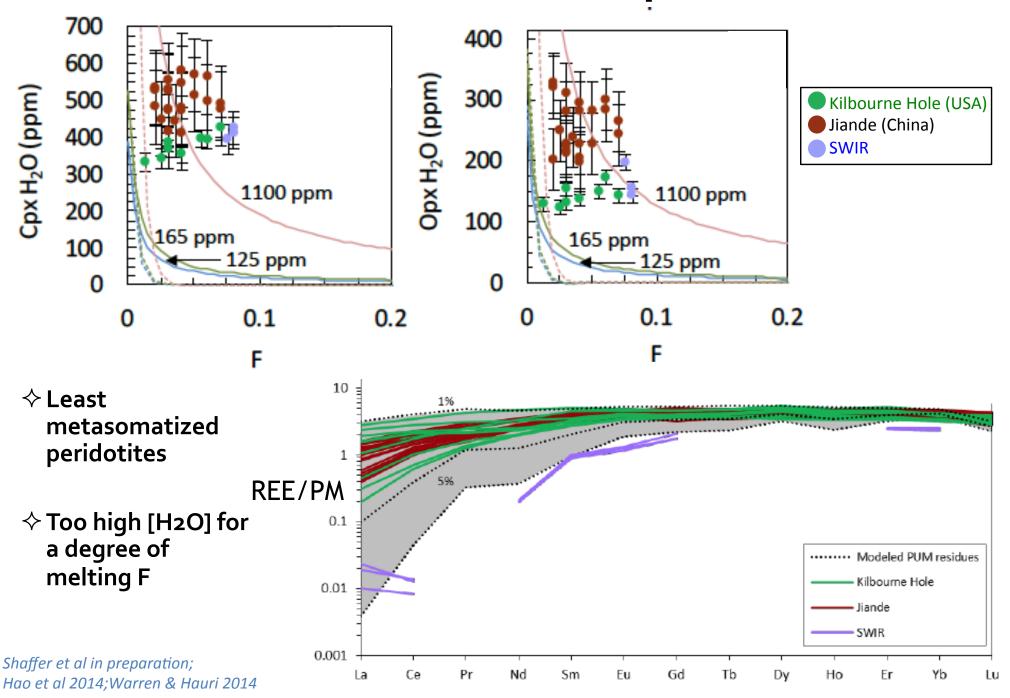


Image Credit: Demouchy & Bolfan-Casanova 2016

- ➤ Similar [H<sub>2</sub>O] for all tectonic settings
- > Subduction peridotite not more wet (except CO Plateau)



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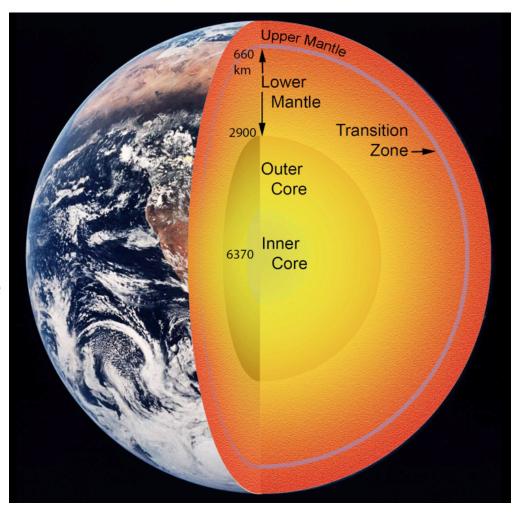
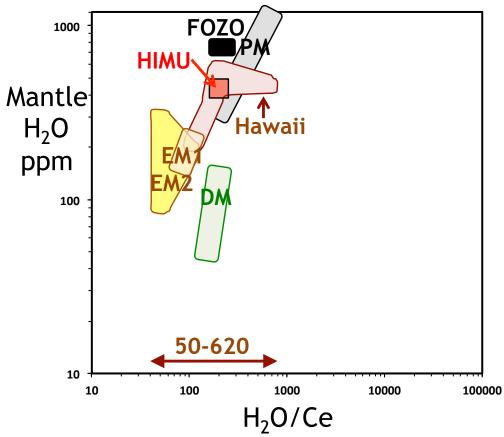


Image Credit: Japan Times

### Water in the oceanic mantle



- Water in melt inclusions & undegassed glasses
- Oceanic island basalts (OIB) have more water than MORB
- Oceanic basalts: water & Ce similarly incompatible

McDonough & Sun 1995; Palme & O'Neill 2007; Dixon et al 2001-08; Salters & Stracke 2004; Cabral et al. 2014; Workman et al. 2004-06; Kendrick et al. 2014-15; Jackson et al. 2007, 2015; Métrich et al. 2014; Seaman et al. 2004

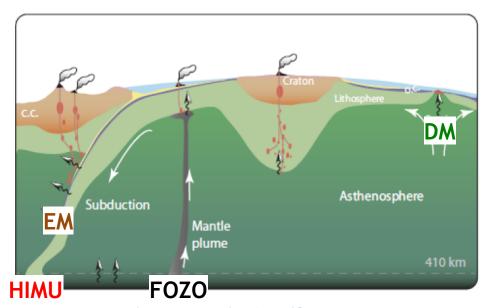
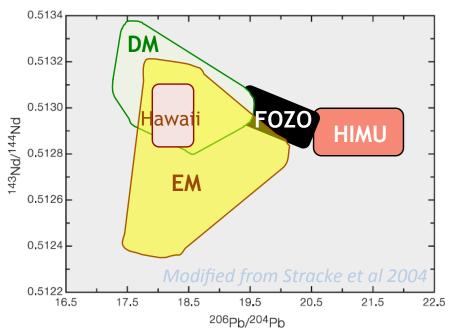
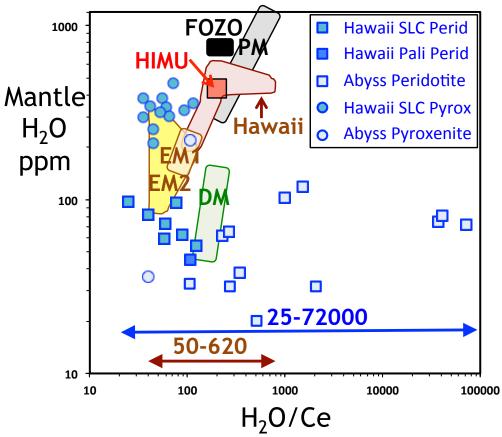


Image Credit: Demouchy & Bolfan-Casanova 2016



#### Water in the oceanic mantle



- Oceanic basalts: water & Ce similarly incompatible
- Oceanic peridotites: water & Ce decoupled
- → Large scale re-equilibration of water?
- → Partition coefficient problem?

McDonough & Sun 1995; Palme & O'Neill 2007; Dixon et al 2001-08; Salters & Stracke 2004; Cabral et al. 2014; Workman et al. 2004-06; Kendrick et al. 2014-15; Wallace 20024; Jackson et al. 2007, 2015; Métrich et al. 2014; Seaman et al. 2004; Peslier & Bizimis, 2015; Bizimis & Peslier 2015; Warren & Hauri, 2014; Peslier et al unpublished

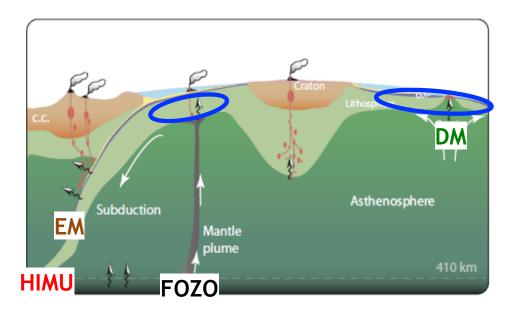
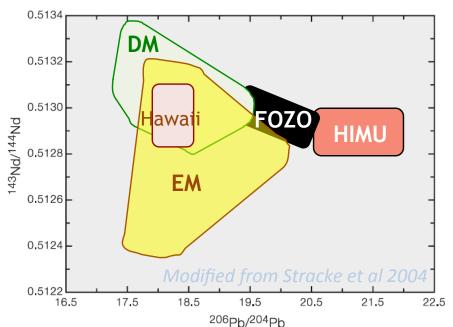
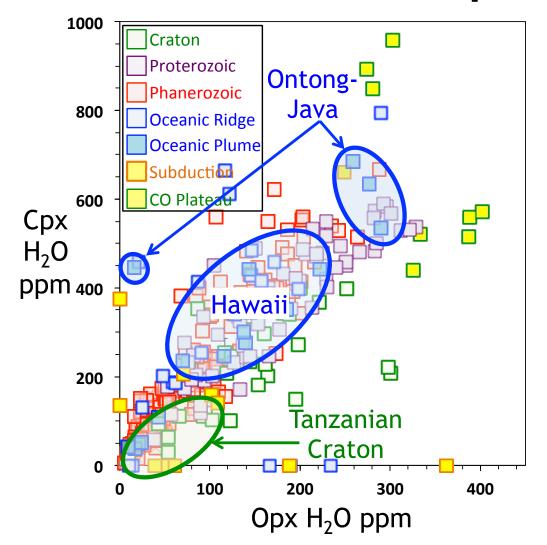


Image Credit: Demouchy & Bolfan-Casanova 2016



## Plume-lithosphere interaction



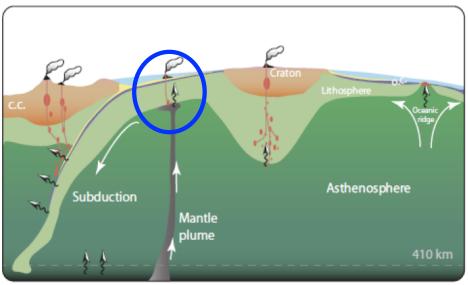


Image Credit: Demouchy & Bolfan-Casanova 2016

- ➤ Similar [H<sub>2</sub>O] for all tectonic settings
- Subduction peridotite not more wet (except CO Plateau)
- Plume interaction (Hawaii, Ontong-Java, Tanz craton) lithosphere not more wet

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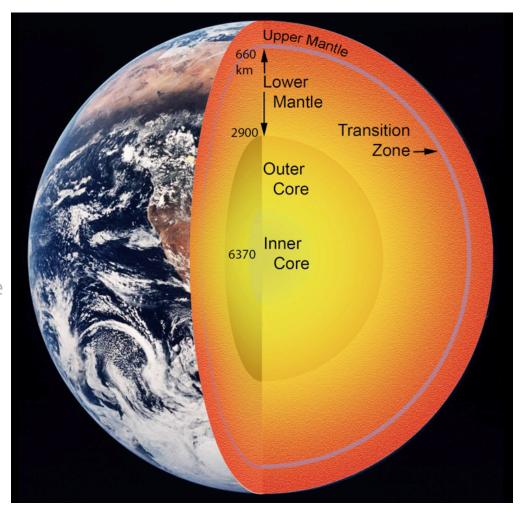


Image Credit: Japan Times

INTERNATIONAL GEOLOGY REVIEW http://dx.doi.org/10.1080/00206814.2015.1056758



#### **REVIEW ARTICLE**

#### Diamonds and water in the deep Earth: a new scenario

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#### ABSTRACT

Earth is a water planet, but how much water exists on and in the Earth? Is the water limited to the Earth's surface and limited depths of our planet (molecular water of the hydrosphere), or do deep reservoirs of hydrogen and oxygen really exist as proposed in recent works but not yet proven? Due to the importance of H<sub>2</sub>O for life and geological processes on the Earth, these questions are among the most significant in all of the Earth sciences. Water must be present in the deep Earth as plate tectonics could not work without water as a major driving force that lowers both viscosity and density of the solid mineral phases of the interior and controls the onset of melting. On subduction, water is returned to the hydrosphere first by dewatering of hydrous phases and second by melting and arc magmatism in and above the subducting slab. The mantle is composed of oxygen minerals, and the extent to which hydrogen is dissolved in them constitutes the true reservoir of the planet's water. Are 'deep water and diamonds' intimately related as indicated in the title of the present article? What is the connection between these two important terrestrial materials? The necessity to review this issue arises from the recent discovery of a strongly hydrous ringwoodite in a Brazilian diamond. As ringwoodite constitutes 60% or more of the lower part of the transition zone, between 525 and 660 km depth, this could correspond to a huge amount of water in this region, comparable or greater in mass to all of Earth's hydrosphere. If the water found in this ringwoodite is representative of the water concentrations of the transition zone, then estimates of Earth's total water reservoir are in need of major revision. This work is an attempt at such a revision.

#### ARTICLE HISTORY

Received 5 April 2015 Accepted 27 May 2015

#### KEYWORDS

diamond; water; Earth; reservoir; mantle

spe500-13 2nd pgs page 431

The Geological Society of America Special Paper 500 2013



#### Whole Earth geohydrologic cycle, from the clouds to the core: The distribution of water in the dynamic Earth system

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Icarus 248 (2015) 89-108



Contents lists available at ScienceDirect

#### Icarus

journal homepage: www.elsevier.com/locate/icarus



Chemical Geology 418 (2015) 6-15

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Accretion and differentiation of the terrestrial planets with implications for the compositions of early-formed Solar System bodies and accretion of water



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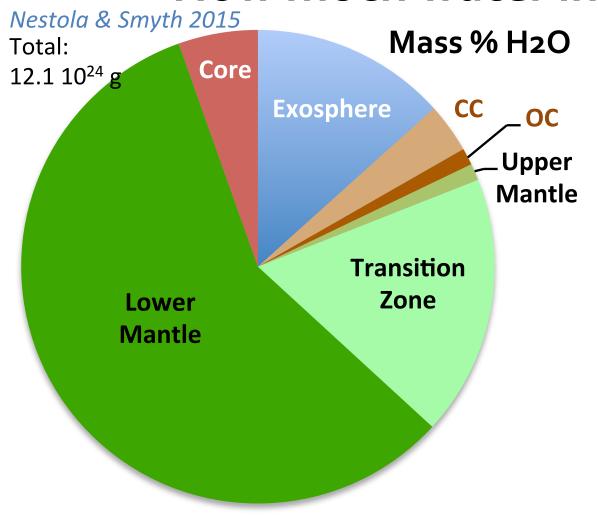
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Hydrous minerals and the storage of water in the deep mantle



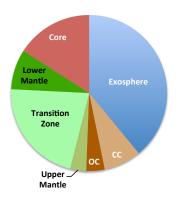
Department of Earth Science, Graduate School of Science, Tohoku University, Sendai, 980-8578, Japan V.S. Sobolev Institute of Geology and Mineralogy, Siberian Branch, Russian Academy of Sciences, Novosibirsk, 630090, Russia



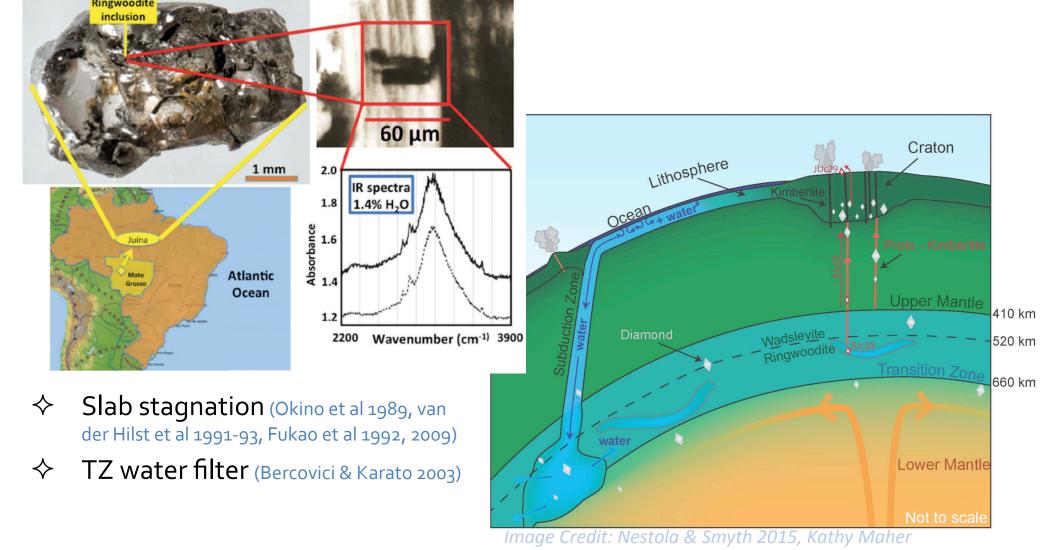


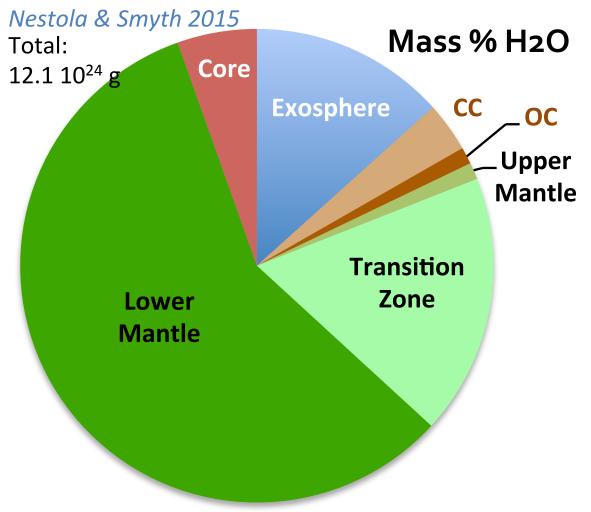
Bodnar et al 2013

Total: 3.61 10<sup>24</sup> g



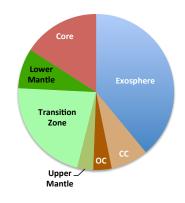
- ♦ Transition zone: probably wet
- High [H2O] in ringwoodite, wadsleyite and majorite from exp (Inoue, Smyth, Ohtani, Demouchy, Bolfan-Casanova)
- Phase Egg (AlSiO<sub>3</sub>(OH)) & hydrous ringwoodite as diamond inclusions (Wirth et al 2007; Pearson et al 2014)
- Consistent with electrical conductivity & seismic tomography (Karato 2001, Yoshino 2010, Koyama et al 2006)





Bodnar et al 2013

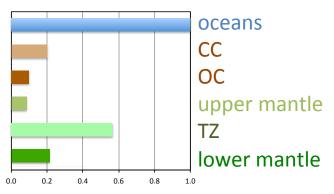
Total: 3.61 10<sup>24</sup> g



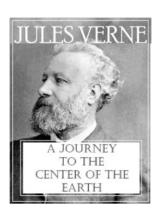
- ♦ TZ: uncertainty on [H2O] of ringwoodite, wadsleyite and majorite.
  - Phase Egg & hydrous ringwoodite as diamond inclusions (Wirth et al 07, Pearson et al 14)
- ♦ Back to problem of storage capacity (exp) vs actual water content
- $\diamond$  Uncertainty on amount of hydrous phases (B, Egg, D, H,  $\delta$ -AlOOH) (Ohtani, Ghosh, Walter)

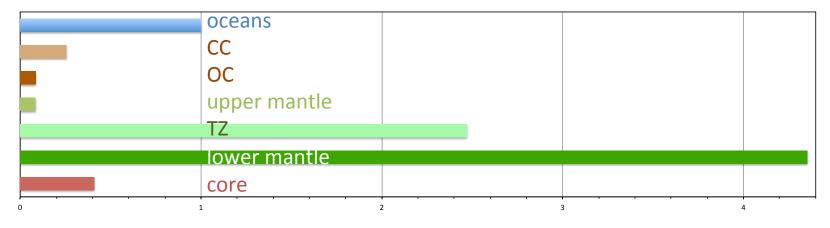
### How much water in the Earth?

#### X mass oceans



Bodnar et al 2013





Nestola & Smyth 2015



Rubie et al 2015
From accretion models

## Earth total water content

- ♦ Type of data
- ♦ Water in the mantle
  - ♦ Definition
  - ♦ Importance
  - ♦ Distribution & controls
    - ♦ H diffusion
    - ♦ Water in the continental mantle lithosphere
    - ♦ Water in the lithosphere
    - ♦ Water in the oceanic lithosphere
- ♦ The big picture
  - ♦ Water in the Earth layers
  - **♦ Fluxes**
- ♦ Comparison with other differentiated planetary bodies

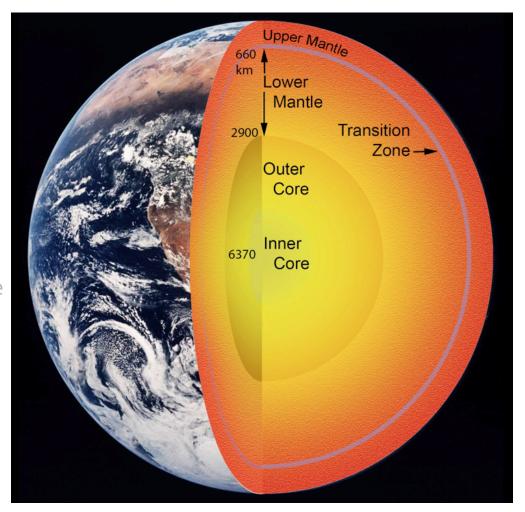


Image Credit: Japan Times

Fluxes in 10<sup>11</sup> kg/year

Oceans:  $1.4x10^{21}$  kg  $10^{11}$  kg = 7 ppb oceans

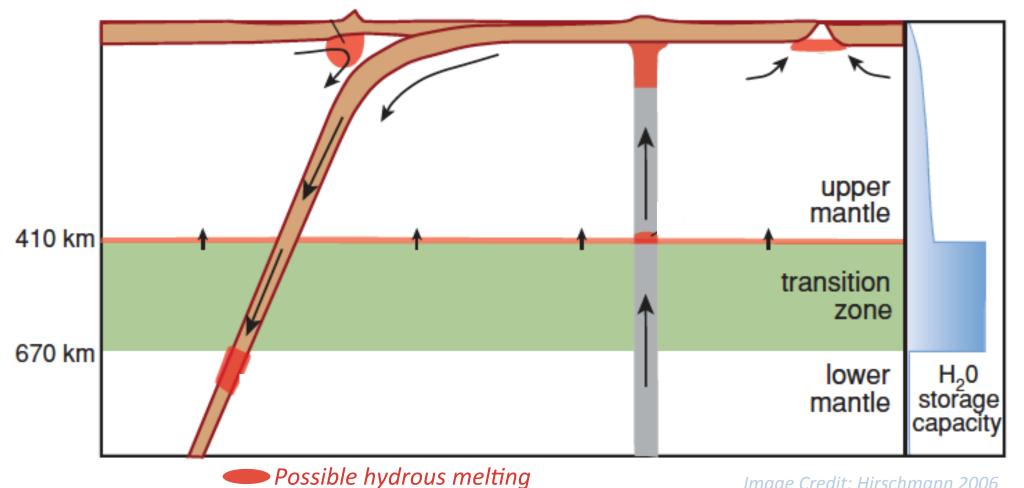
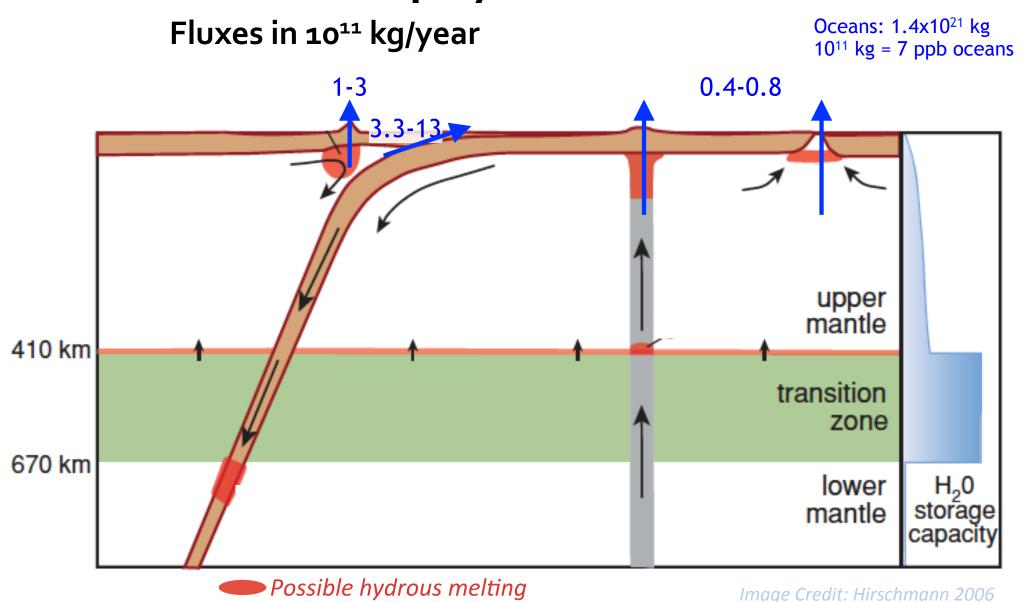
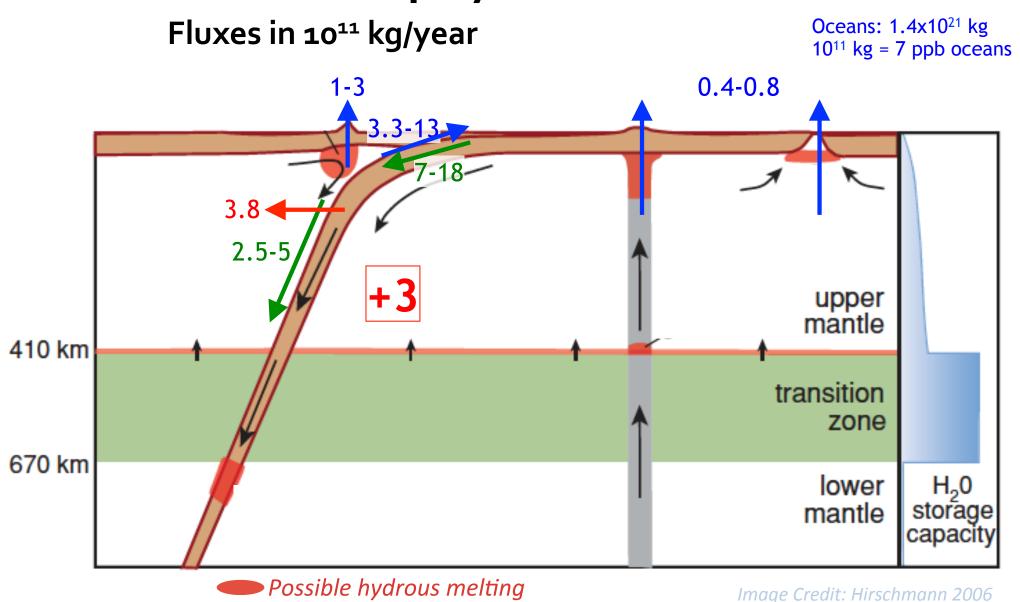
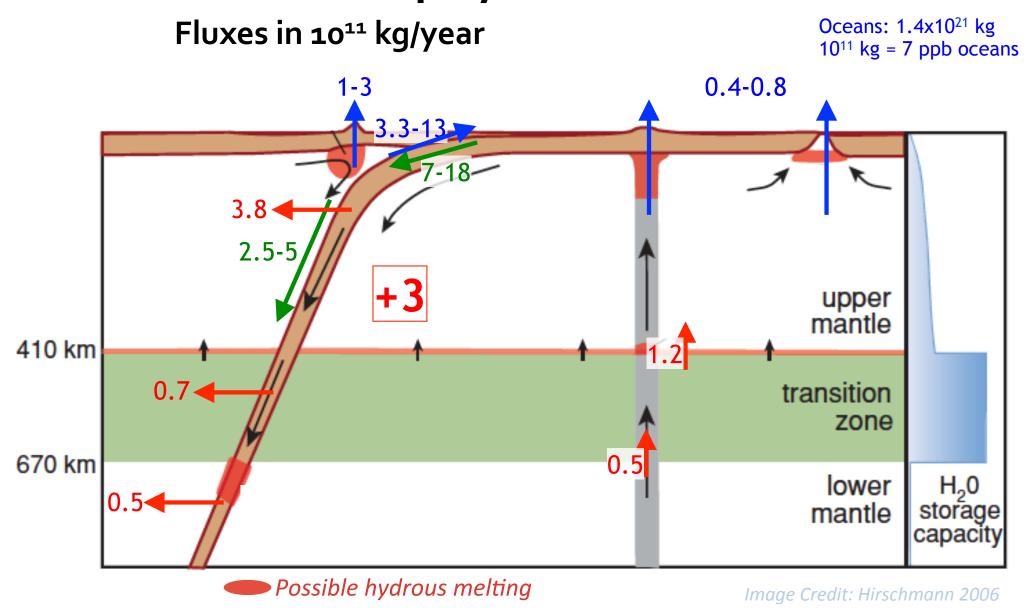
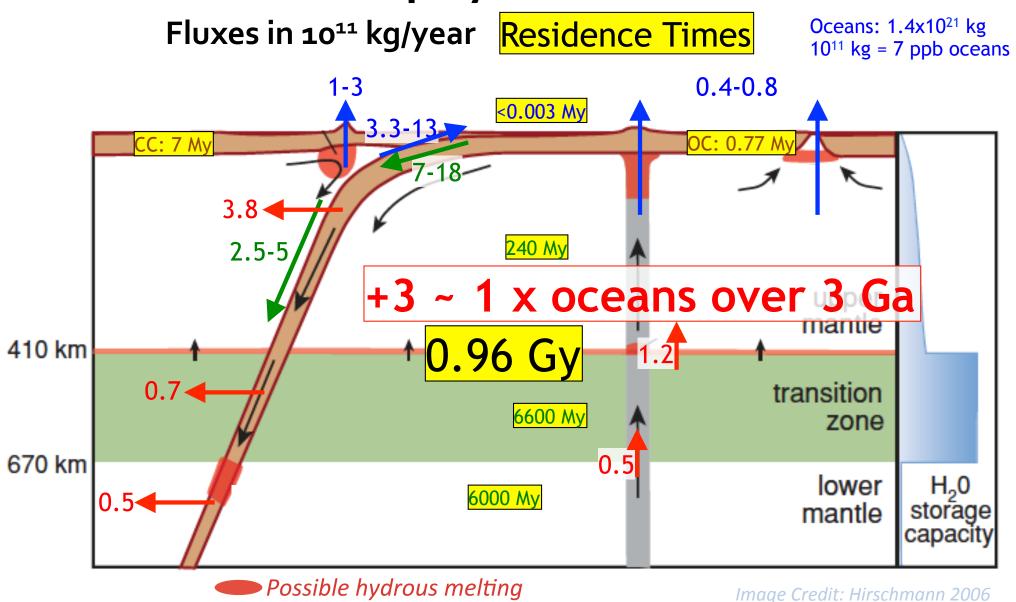


Image Credit: Hirschmann 2006









# Earth total water content

- ♦ Type of data
- ♦ Water in the mantle
  - ♦ Definition
  - ♦ Importance
  - ♦ Distribution & controls
    - ♦ H diffusion
    - ♦ Water in the continental mantle lithosphere
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  - ♦ Water in the Earth layers
  - ♦ Fluxes
- ♦ Comparison with other differentiated planetary bodies

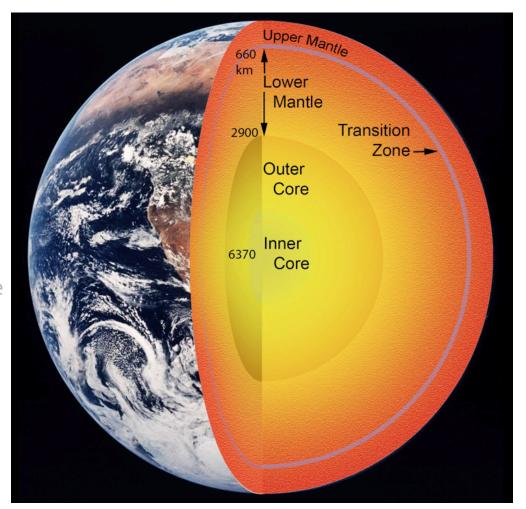
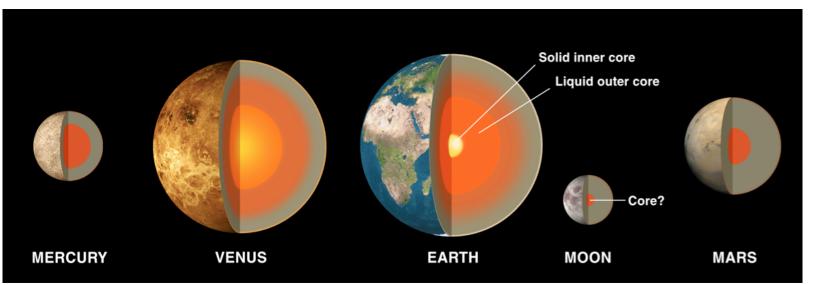
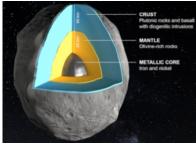


Image Credit: Japan Times

# Water in the inner solar system





**VESTA** 

### Upper mantle ppm H<sub>2</sub>O from samples

**150-350** 

5-100

**130-230**(Ch)

>0?

**160**(LMO) **15-47**(Sh)

Mantle water from accretion models (Mantle H<sub>2</sub>O-Core H)

**961**-*35* 

948-33

0-4769-4

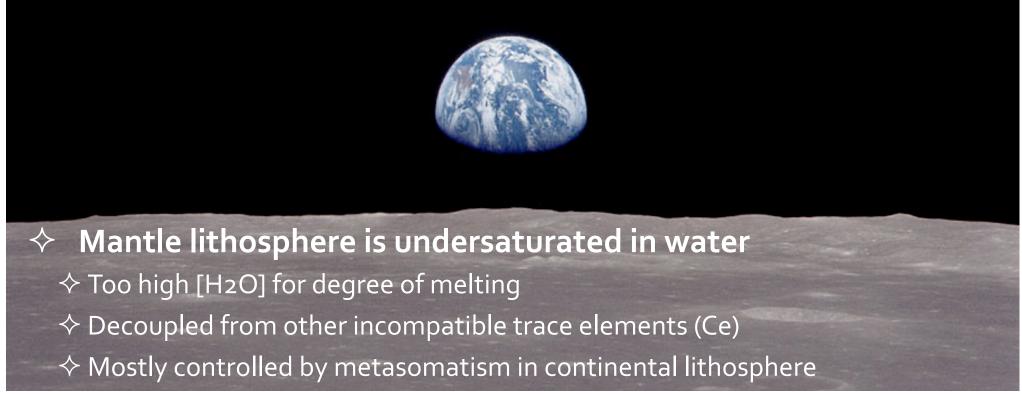
Origin

CC

CC

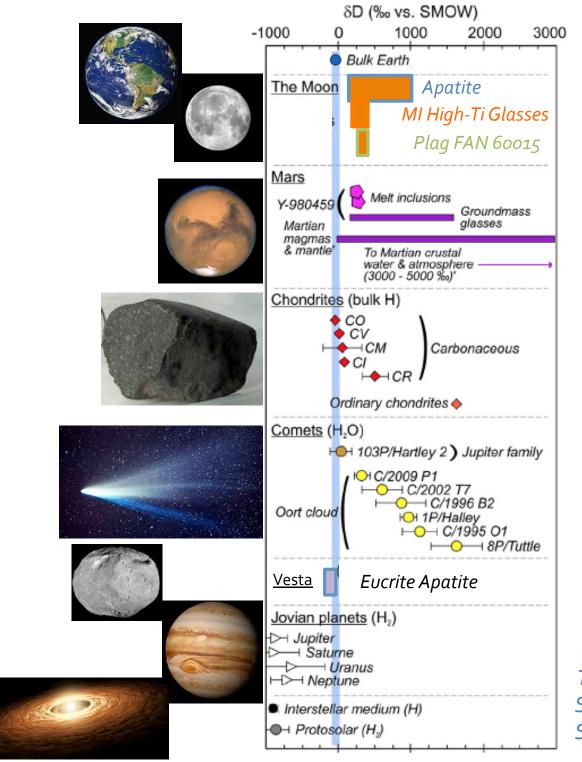
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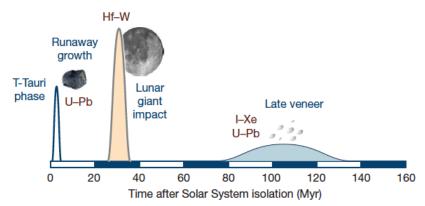
### **Conclusions**

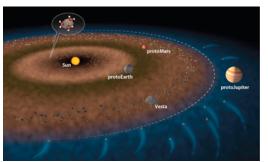


- ♦ All tectonic settings have similar [H2O]
- ♦ The transition zone is rich in water
- ♦ Water content of lower mantle and core?
- ♦ Water deep cycle

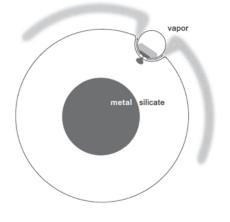
  - ♦ Residence time of water in Earth: ~ 1 Ga







#### or/and



Tartèse et al. 2013; Greenwood et al. 2011, Saal et al., 2013; Albarède et al., 2009, 2013; Sarafian et al., 2013-2014; Hui et al., 2015

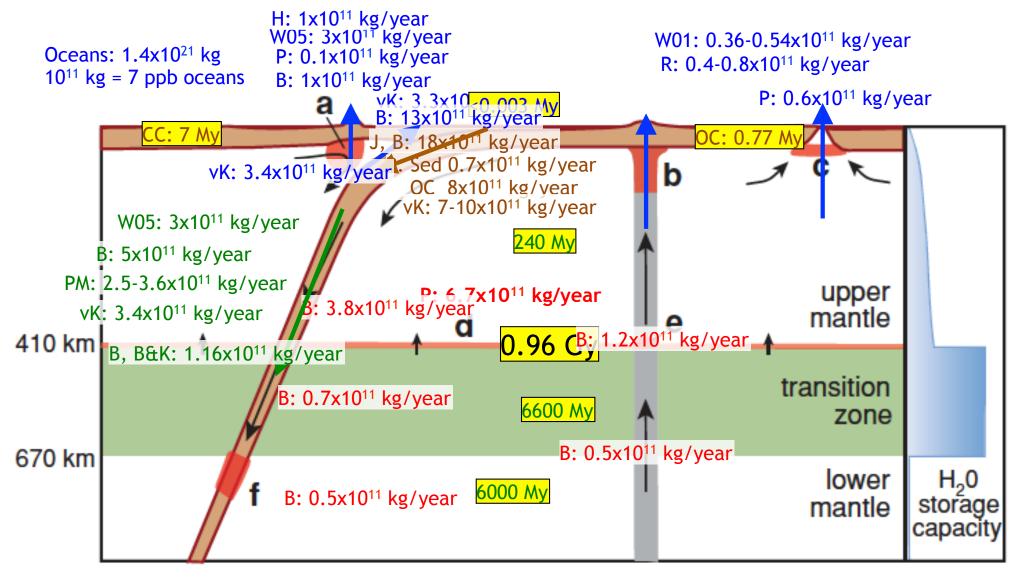
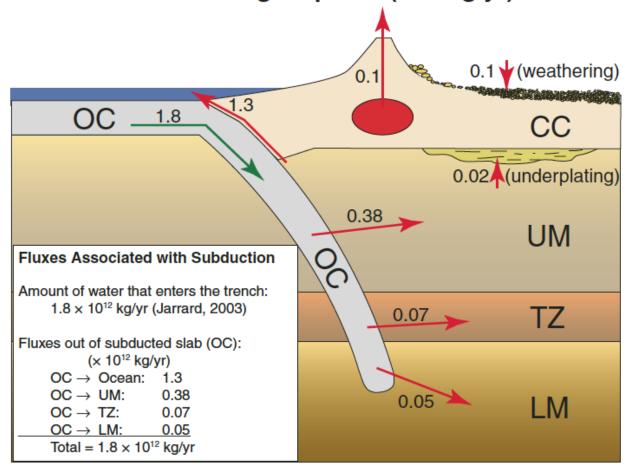
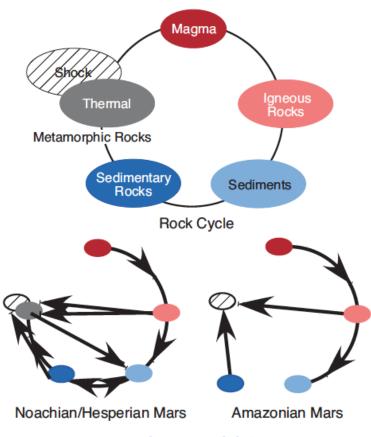


Image Credit: Hirschmann 2006

#### Fluxes in geosphere (10<sup>12</sup> kg/yr)



Bodnar et al 2013



McSween 2015

Parai & Mukhopadhyay 2012